Mon. Apr. 02, 2018

- Satellite Gravity Measurements (GRACE, GRAIL) – Use slides posted for Friday
- Linear (Spectral) Mixing

- Reading: Chapter 9 ("Environmental" Remote Sensing")
	- Once again -- Satellites old but principles still apply

"Linear Spectral Unmixing 1"

- Linear: Assume resulting spectra is weighted average of endmember spectra.
	- (Only true for macroscopic mixing. Microscopic (intimate) mixing is often non-linear.
- Following is true for every pixel. Repeat mathematics for each pixel.
- Observe some spectrum $I(\lambda)$
- There are N possible minerals present, each occupying a fraction f_j of the pixel (for $j=1$, N) – Usually you assume \angle J j i.e. pixel is fully occupied $\sum_{j\in J,N} f_j = 1$ $j = 1, N$
- You know the spectrum $R_i(\lambda)$ for each of the j=1,N minerals
- For linear mixing

$$
I(\lambda) = \sum_{j=1,N} f_j R_j(\lambda)
$$

- Given $I(\lambda)$ and $R(\lambda)$, find the fractions f_i
- *Complications*
	- *Noise*
	- *Observations only at a limited number of*

Mineral Spectra: Example of linear mixing

"Linear Spectral Unmixing 2"

• Simple example: Single wavelength, and observed $I_1 = 0.5$ Two minerals, with $R_{11} = 0.4$ and $R_{12} = 0.7$ (I'll explain the first R subscript shortly)

$$
I_1 = R_{11} f_1 + R_{12} f_2
$$

= R_{11} f_1 + R_{12} (1 - f_1)
0.5 = 0.4 f_1 + 0.7 (1 - f_1)
0.5 = -0.3 f_1 + 0.7
-0.3 f_1 = -0.2
f_1 = 0.67 and f_2 = 1 - 0.67 = 0.335

• However if we do not have endmember minerals whose reflectances bracket the observed value, you get "nonsense" results with $f_i < 0$ or $f_i > 1$

Single wavelength, and observed $I_1 = 0.5$ Two minerals, with $R_{11} = 0.4$ and $R_{12} = 0.3$

|

$$
0.5 = 0.4 f1 + 0.3 (1 - f1)
$$

\n
$$
0.5 = 0.1f1 + 0.3
$$

\n
$$
0.1f1 = 0.2
$$

\n
$$
f1 = 2.0 and f1 = 1 - 2.0 = -1.0
$$

- More sophisticated versions of "unmixing" can enforce constraints $0 \le f_i \le 1$ at expense of "imperfect" fits to the observed reflectance.
- Other versions relax requirement that $f_1+f_2=1$ and (when you have many wavelengths and pixels) allow presence of an unknown component whose spectrum is similar to the <u>average</u> spectrum of all the pixels.

"Linear Spectral Unmixing 3" • For linear mixing *^I*(*^λ*)= ∑ *j*=1, *N* $f_j R_j(\lambda)$

• If you only have data at discrete wavelength λ_i for i=1,M

$$
I(\lambda_i) = \sum_{j=1,N} f_j R_j(\lambda_i) \quad \text{or,defining } I_i \equiv I(\lambda_i) \text{ and } R_{i,j} \equiv R_j(\lambda_i)
$$

$$
I_i = \sum_{j=1,N} f_j R_{i,j} = \sum_{j=1,N} R_{i,j} f_j
$$

• Which can be written in matrix form:

 $\vert \cdot \vert$ I_{1} I_{2} \vdots I_M^{\cdot} | I = $\int_{\mathbf{n}}$ R_{11} R_{12} … R_{1N} R_{21} R_{22} … R_{2N} \vdots \vdots \vdots R_{M1} R_{M2} … R_{MN} $\left| \left| \dot{f}_{N} \right| \right|$ *f* 1 *f* 2 \vdots $\dot{\left|f_{N}\right|}$

- Each vertical column in the matrix is the spectrum of a given mineral.
- We want to determine f_i given the known I_j and R_{ij} We have M simultaneous equations (= # λ_i) with N unknowns $(=\# \text{ of minerals})$
- Including the constraint gives us one more equation.

$$
\sum_{j=1,N} f_j = 1
$$

- Since the number of unknowns can at most equal the number of equations, if we have M wavelength bands we can model at most $N = M+1$ minerals.
- If we have more wavelengths than necessary, the problem is "overdetermined" but we can then produce best fits which in some "average" sense match the observed spectrum "as well as possible".

Linear Mixing Application to Io

Loki: Most Active Volcanic Region on Io

Loki "Bergs": Fumaroles?

- Thought to be large $(200 \text{ km diameter})$ lava lake
- Bright spots may be fumarole deposits
- We're trying to understand their composition

Colors of the "bergs"

 1.0

Reflectance

 $0.0 - 3$

- The "bergs" almost disappear in the violet (The largest are still slightly visible)
- Consistent with a sulfur composition
- Not consistent with (primarily) SO_2 composition
- Basalt Also makes sense in terms of temperatures on patera

Plot Violet Brightness vs. Blue Brightness

- Imaged reduced and projected to simple cylindrical maps using ISIS3
- Brightnesses are "I/F" values, \bullet returned by standard ISIS3 code.
	- While photometric corrections have not been $\overline{}$ applied, phase angles are relatively small (and similar) and most photometric corrections at these angles are relatively small
	- Linear mixing between endmembers in a twobrightness plot like this simply produces mixing "lines" connecting the endmember reflectance points.

Violet vs. Blue Brightness: SW (dark) Patera Only

Violet vs. Blue Brightness: SW and S Patera

Violet vs. Blue Brightness: SW & S Patera + Bergs

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Composition of other Components at Loki?

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- 1) While the very dark SW corner is likely to be bare basalt -- the slight vertical (V) jump in reflectance between it and the rest of the patera suggests something more than just sulfur deposition affects the violet color. Perhaps this is an age-related "weathering" effect.
- 2) The SW Bay and the "overflow region can be explained by basalt, covered by increasing amounts of sulfur.
- 3) The SW Overflow "Bathtub Ring" and the island COULD be explained by a mixture of more reflective small-grain S -- combined with a dark component like basalt -- but a more natural explanation is:
- 4) Once coverage by course sulfur becomes essentially 100%, the thermal environment may allow condensation of small amounts of SO_2 , which then begin to raise the Violet albedo significantly.

Determining "Endmembers" using 2-D Scatter Plots

•Compute PCA or MNP components •Plot one band of output vs. another •Possible endmembers are shown by colors in following

