



1 Crustal structure and thickness along the Yellowstone hot 2 spot track: Evidence for lower crustal outflow from beneath 3 the eastern Snake River Plain

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9
10 [1] Receiver functions from seismic stations about the Yellowstone hot spot track are migrated to depth
11 using a V_p/V_s map constructed from stacking of the direct and free surface Moho reverberations (i.e.,
12 H-K analysis) and a shear velocity tomogram constructed from surface wave measurements. The thickest
13 crust (48–54 km) resides in the Wyoming province beneath the sampled Laramide age blocks, and the thin-
14 nest crust (32–37 km) resides in the Montana Basin and Range province. The eastern Snake River Plain
15 (ESRP) crust is thickest (47 km) at its NE end beneath the young calderas and thinnest (40 km) at its
16 SW end beneath the older Twin Falls caldera. Two ESRP crustal thickness domains are found: (1) at
17 the older Twin Falls and Picabo calderas, the mean ESRP crust is 4 km thicker with respect to its margins
18 and (2) adjacent to the Heise caldera field, the mean ESRP crust is 4 km thicker with respect to its SE margin
19 crust but no thicker with respect to its NW margin crust. This lobe of anomalously thick crust is explained
20 as resulting from lower crustal outflow from beneath the Heise caldera field. Confirmation of these crustal
21 thickness variations is provided by inspection of common conversion point (CCP) stacks that delineate sev-
22 eral secondary features: the top of a thick high-velocity (3.9 km/s) lower crust layer within the Wyoming
23 province up to 17 km thick and a paired negative and positive amplitude arrival at 12 km depth and 18 km
24 depth beneath the Yellowstone Caldera. This paired arrival would be consistent with a low-velocity zone
25 perhaps associated with magma staging beneath the caldera. Our most important finding is that the mag-
26 matic loads injected into the ESRP crust over the last 4–12 Myr, in tandem with the ESRP crustal viscosity
27 structure, have been sufficient to drive significant outflow of the ESRP lower crust.

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35 1. Introduction

36 [2] The Yellowstone hot spot track manifests a
37 sequence of calderas that begin 16.9 Myr ago near
38 the tristate region of Oregon-Idaho-Nevada with
39 the calderas propagating to the NE (Figure 1). The
40 sequence of calderas trend to the NE from the
41 Bruneau-Jarbridge (12.7–10.5 Ma) to the Twin
42 Falls (10.5–8.6 Ma) to the Picabo (10.2–9.2 Ma) to
43 the Heise (6.6–4.4 Ma) caldera fields. All together,
44 these calderas reside within the structural down-
45 warp termed the eastern Snake River Plain (ESRP)
46 [Perkins and Nash, 2002; Bonnicksen et al., 2007;
47 Anders, 2009; Rodgers and McCurry, 2009;
48 Leeman et al., 2008]. The most recent 2.1–0.6 Ma
49 Huckleberry Ridge, Island Park, and Yellowstone
50 calderas reside primarily upon the Yellowstone
51 Plateau. Explanation of the ESRP downwarp
52 requires knowledge of the crustal magma injection
53 volumes and composition, time-integrated exten-
54 sion, and time-integrated lower crustal flow fluxes
55 [McQuarrie and Rodgers, 1998; Stachnik et al.,
56 2008; Rodgers and McCurry, 2009]. The largest
57 uncertainty with respect to calculating an ESRP
58 crustal mass balance is knowledge of the pre-hot
59 spot magmatism and extension and the crustal
60 thickness and density. The pre-hot spot conditions
61 can only be roughly estimated from the tectonic
62 history of this region [Hamilton, 1989; Dickinson,
63 2006; Foster et al., 2006]. However, the modern
64 day crustal thickness and density structure can be
65 estimated from the combination of the Earthscope
66 Transportable array and previous PASSCAL seis-
67 mic data.

68 [3] Previous petrologic and geochemical modeling
69 suggests that 8–14 km of basaltic magma was in-
70 jected primarily into the 8–18 km depth range be-
71 neath the individual calderas [Bonnicksen et al.,
72 2007; Hanan et al., 2008; McCurry and Rodger,
73 2009; Leeman et al., 2008]. Heat budget modeling
74 of the duration and volumes of the Rhyolite eruption
75 volumes also requires about 10 km of basaltic
76 magma injection into the mid to upper crust [Leeman
77 et al., 2008]. The midcrust is the preferred Mixing-
78 Assimilation-Hybridization (MASH) region where
79 the Rhyolitic liquids are distilled via fractionation
80 and modest levels of crustal assimilation [Hildreth
81 and Moorbath, 1988]. Geophysical evidence for
82 the fossil (crystallized and cooled) midcrustal
83 MASH zone beneath the ESRP derives from
84 gravity analysis [Sparlin et al., 1982; Lowry and
85 Smith, 1995; Lowry et al., 2000; DeNosaquo et
86 al., 2010], seismic refraction analysis [Smith et
87 al., 1982; Sparlin et al., 1982], and local earth-

quake analysis [DeNosaquo et al., 2010]. In ad- 88
dition, our surface wave seismic tomogram finds 89
an approximately 10 km thick high-velocity layer 90
in the ESRP midcrust with the top of the high- 91
velocity layer at 15–25 km depth [Stachnik et al., 92
2008]. 93

[4] As the magma injections beneath the calderas 94
cool on 1–2 Myr time scales [Anders and Sleep, 95
1992], the midcrustal sill complex (MCS) is pre- 96
dicted to become denser than the surrounding country 97
rocks [McCurry and Rodger, 2009; DeNosaquo et 98
al., 2010]. This high-density MCS would thus 99
become a positive load on the lower crust which 100
could force the lower crust to flow outward on 101
million year time scales if the crustal viscosity is 102
sufficiently low [Buck, 1991; McQuarrie and 103
Rodgers, 1998; Royden et al., 2008]. The ESRP 104
can thus be viewed as a magmatic time machine 105
that loads the lower crust; this factor permits assess- 106
ment of potential lower crustal flow if accurate 107
crustal thickness maps are available. 108

2. Data and Methods 109

[5] The broadband seismic recording analyzed 110
(Figure 1) derive from Earthscope Transportable 111
array data, six three-component short-period Uni- 112
versity of Utah Seismic Network seismometers in 113
Yellowstone Park and five dominantly broadband 114
PASSCAL seismic experiments: the 1993 eastern 115
Snake River Plain line array [Saltzer and Humphreys, 116
1997], the N–S and NW–SE Deep Probe passive line 117
arrays [Dueker and Yuan, 2004], the 2000–2001 118
Yellowstone array [Yuan and Dueker, 2005], the 119
1999–2000 Billings array [Yuan et al., 2008] and the 120
2000–2001 Stanford Snake River Plain axis array 121
[Walker et al., 2004]. Teleseismic P wave arrivals 122
for receiver function analysis were windowed from 123
the continuous data for all events with body wave 124
magnitudes >5.3 and the seismic recording compo- 125
nents rotated into the vertical, radial, tangential co- 126
ordinate system. To source normalize the events, a 127
multitaper spectral correlation method was used 128
[Park and Levin, 2000; Helffrich, 2006]. The verti- 129
cal component was used as an estimate of the source 130
function and deconvolved with a dynamic water 131
level [Clayton and Wiggins, 1976] derived from 132
the preevent noise spectral amplitude. The result- 133
ing receiver functions were then culled of noisy 134
traces by removing the 20% of the radial component 135
receiver function with RMS amplitudes greater than 136
three times the mean of the data set. In addition, a 137
visual inspection was done to remove dead and/or 138
harmonic traces. Finally, all the radial receiver 139

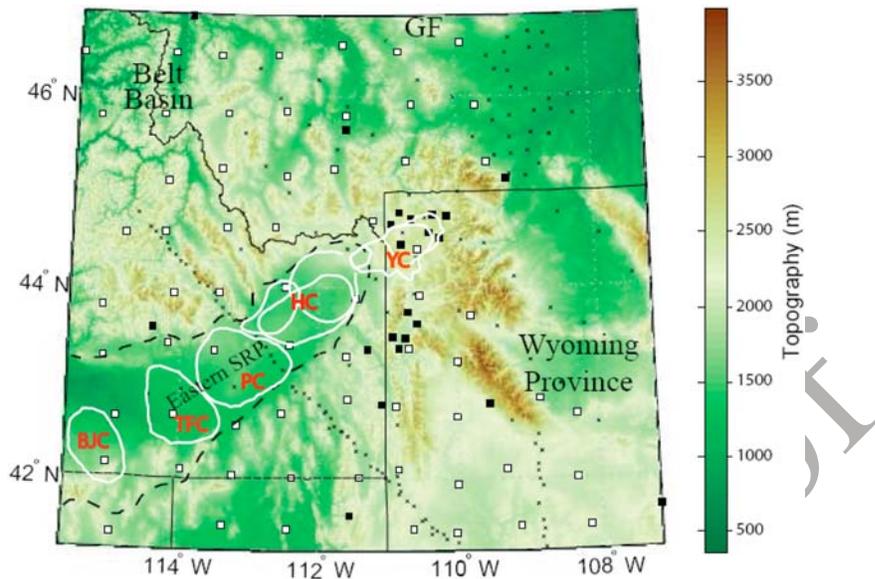


Figure 1. Seismic stations and topography. The eastern Snake River Plain is outlined with the black dashed line, and the volcanic calderas are denoted as white outlines with labels: 12.7–10.5 Ma Bruno-Jarbidge (BJC), 10–8.6 Ma Twin Fall (TFC), 10.1–10.3 Ma Picabo (PC), the 6.6–4.2 Ma Heise field (HC, a composite of at least three calderas), and the 2.1–0.6 Ma Huckleberry Ridge/Island Park/Yellowstone caldera (YC) field. Seismic stations used in this study are denoted as follows: transportable array stations (white squares), four PASSCAL experiment stations (crosses), and University of Utah/Grand Teton Park/National Seismic network stations (filled squares).

140 functions were linearly stacked with moveout
141 corrections to provide 134 station stack traces
142 (Figure 2).

143 [6] For each of the individual stations, a crustal
144 thickness and bulk crustal V_p/V_s analysis (i.e., H-K
145 analysis) is performed [Zandt et al., 1995; Zhu and
146 Kanamori, 2000] to constrain the mean crustal
147 thickness and V_p/V_s value in the seismic sampling
148 cone beneath each station (Figure 3). The weight-
149 ing of S wave arrivals converted at the Moho in the
150 H-K stacks is 0.6 (direct arrival), 0.3 (2p1s free
151 surface reverberation), and 0.1 (2s1p free surface
152 reverberation). The shear velocity at each station
153 was specified to be the 1-D velocity profiles which
154 were extracted at each station point (Figure 4) from
155 the shear wave velocity tomogram [Stachnik et al.,
156 2008]. The V_p/V_s and crustal thickness marginal
157 probability density functions were estimated using
158 bootstrapping with replacement [Efron and
159 Tibshirani, 1986]. The probability functions de-
160 rived from the bootstrapping are generally peaked
161 unimodal functions (Figures 3b and 3c): indicating
162 reasonable resolution of the trade-off between
163 crustal thickness and V_p/V_s variations. The best
164 estimate of the two model parameters was consid-
165 ered to be the mode of the distributions. The pa-
166 rameter errors were estimated from these probability

functions by estimating the standard deviation about
the mode of the probability function.

[7] To construct maps of the V_p/V_s (Figure 5),
crustal thickness and crustal thickness errors
(Figure 6), a least squares spline algorithm with a
second derivative regularization term was used to
interpolate the single-station results. The RMS
difference between the spline predicted parameter
values and the H-K measured crustal thickness and
 V_p/V_s values were small due to the spatial coherence
of the single-station measurements. The crustal
thickness errors (standard deviation) from boot-
strapping the peak arrival time of the Moho arrival
are <1.5 km (Figure 6b). However, this crustal
thickness error estimate does not include the
migration velocity model uncertainties associated
with the shear wave velocity model (Figure 4) and
the V_p/V_s map uncertainties used to migrate the
receiver functions (Figure 5). Sensitivity analysis
finds that a 0.2 km/s variation in the bulk shear
velocity would produce a 0.5 km variation in Moho
depth [Zhu and Kanamori, 2000]. Our surface
wave based shear velocity tomogram is a well-
resolved image and has errors <0.2 km/s for the
mean crustal velocity [Stachnik et al., 2008]; thus,
the crustal thickness errors associated with our shear
velocity tomogram migration velocities is estimated

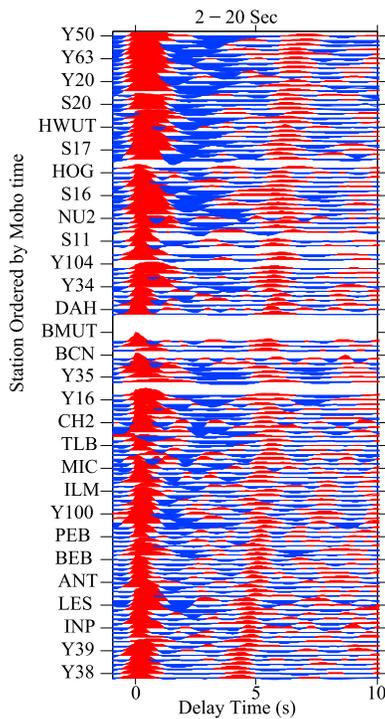


Figure 2. Station mean radial receiver functions for selected stations filtered at 2–20 s band pass. The receiver functions for each station were moveout corrected to 0.06 s/km ray parameter and linearly stacked. The stations are ordered by arrival time of the Moho arrival at 4.3–6.9 s.

194 as <1 km. Sensitivity analysis also shows that a
195 0.1 change in V_p/V_s will produce about a 4 km
196 change in Moho depth [Zhu and Kanamori, 2000].
197 Our V_p/V_s error map (Figure A1) finds that the
198 standard deviation of our bootstrapped V_p/V_s mea-
199 surements is <0.03; thus, crustal thickness errors
200 associated with our V_p/V_s map are estimated as
201 <1.3 km. Assuming the three sources of error are
202 independent, the maximum uncertainty in our
203 crustal thickness maps is estimate as <3.7 km,
204 more typically <2.5 km.

205 [8] The common conversion point (CCP) receiver
206 function images [Dueker and Sheehan, 1997] were
207 constructed using a three dimensional pixel pa-
208 rameterization with 2 km thick layers and 40 by
209 40 km wide CCP bins (Figures 8–11). The bin
210 center points were spaced every 10 km so that
211 adjacent bins had 75% data overlap, but no data
212 overlap at three bin offsets. Time was mapped to
213 depth using our shear velocity tomogram [Stachnik
214 et al., 2008] and the V_p/V_s map (Figure 5) for the
215 crust. For the mantle below the Moho, the IASPEI91

velocity model mean upper mantle V_p/V_s value of 216
1.81 was used. 217

3. Results 218

3.1. Single-Station Interpolated Maps 219

[9] As our prior seismic results reported [Stachnik 220
et al., 2008], the mean crustal velocity map gen- 221
erally shows high (>3.6 km/s) mean velocity 222
Wyoming province crust and low (<3.5 km/s) mean 223
velocity Yellowstone caldera crust (Figure 4). The 224
high-velocity Wyoming crust is primarily due to a 225
high-velocity lower crustal layer imaged by the 226
shear wave tomogram (Figure 11). This layer has 227
been previously imaged as the so-called 7.x mag- 228
matic underplate layer (i.e., with a velocity >7.0 km/ 229
s) from active source studies [e.g., Gorman et al., 230
2002]. In addition, the CCP images find a positive 231
arrival from the top of this high-velocity layer where 232
the station density is highest within the Billings, 233
Montana array (Figures 8, 9, and 11). 234

[10] The V_p/V_s map (Figure 5) shows a reasonable 235
range of 1.76–1.86 (ignoring the map edge values). 236
The mean V_p/V_s is 1.81 compared to global esti- 237
mate mean continental crustal value of 1.79 [Zandt 238
and Ammon, 1995; Christensen, 1996]. A typical 239
quality H-K analysis is shown in Figure 3 which 240
finds a V_p/V_s value of 1.79 ± 0.02 . The maximum 241
error in the V_p/V_s map is 0.03 as found by boot- 242
strapping the H-K analysis. The principle anomaly 243
observed in the V_p/V_s map is the relatively high 244
values (>1.84) along the ESRP and normal values 245
within the Yellowstone Park. 246

[11] The crustal thickness map (Figure 6a) gener- 247
ally shows thick crust (54–48 km) within the 248
Wyoming province crust in Wyoming and eastern 249
Montana. Specifically, thick crust is found beneath 250
the Billings, Montana region and the two sampled 251
Laramide age blocks associated with the Wind 252
Rivers and Beartooth Mountains. But, the seismic 253
sampling under the Bighorns Fault block is too 254
sparse to draw any conclusions with respect to its 255
crustal thickness. Three patches of intermediate 256
thickness (40–44) crust are found within Wyoming 257
surrounding the Wind River and Beartooth Laramie 258
blocks. Thin crust (33–37 km) is found within the 259
Montana Basin and Range Province [Zeiler et al., 260
2005], beneath the sampled Idaho Batholith [Kuntz 261
et al., 2005], and to the south of the central ESRP. 262
The Yellowstone Plateau region has crustal thick- 263
ness of 47–52 km; this thick crust primarily man- 264
ifests the Laramide age shortening associated with 265

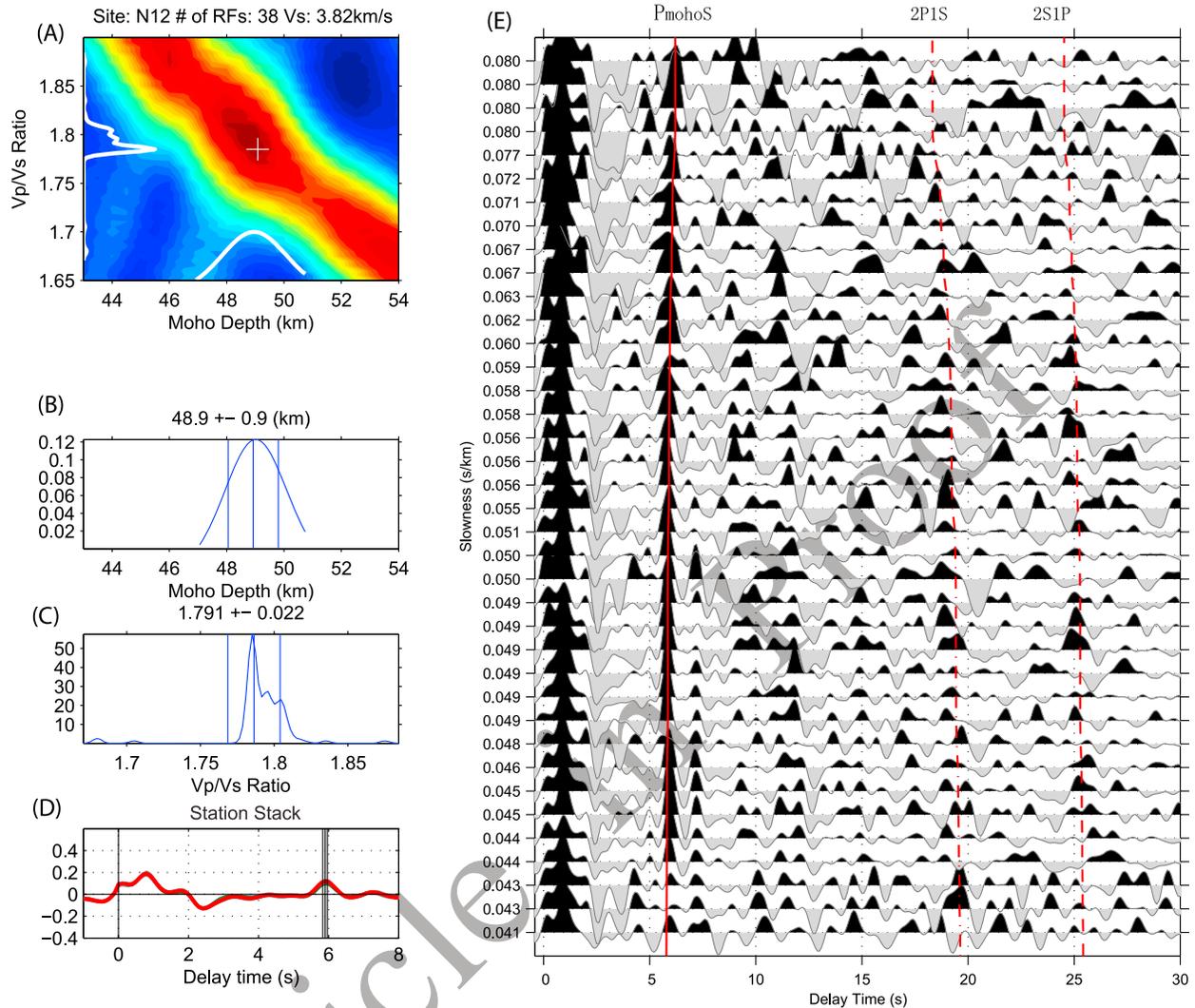


Figure 3. H-K stack analysis for a typical station. (a) H-K stack amplitude for the Moho depth versus V_p/V_s . The one-dimensional probability density function are the white lines along the x and y axes. (b) Estimate of Moho depth (distribution mode) and error as vertical bars. (c) Estimate of crustal V_p/V_s (distribution mode) and error as vertical bars. (d) Mean station moveout corrected radial receiver function. The Moho arrival is marked at 6 s, and amplitude is with respect to the vertical P wave component. (e) Station radial receiver function data with moveout curves for the direct (P_{mohoS}) and two free surface Moho reverberations (2P1S and 2S1P) overlaid.

266 the Beartooth Mountains and the magmatic under-
267 plate that created the high-velocity lower crust be-
268 neath much of the Wyoming Province crust.

269 [12] The ESRP crustal thickness is found to thin by
270 8 km from 49 km at the NE end where the Huck-
271 leberry Ridge/Island Park calderas reside (Figure 7,
272 cross section B) to 41 km beneath the SW end of
273 our ESRP sampling where the Twin Falls caldera
274 resides (Figure 7, cross sections F and G). The most
275 remarkable crustal thickness anomaly observed in
276 the ESRP perpendicular cross sections is a lobe of
277 thicker crust located beneath the NW ESRP margin

adjacent to the Heise caldera field (Figure 7, cross 278
sections C–E). The cross sections through this 279
NW Heise crustal anomaly show 44–46 km crust 280
extends 50–80 km to the NW of the ESRP margin. 281
This NW Heise crustal anomaly contrasts with the 282
cross sections through the older Picabo and Twin 283
Fall calderas (cross sections F and G) where crustal 284
thicknesses >38 km are confined within the ESRP 285
physiographic margins. These reported ESRP 286
crustal thickness values are generally consistent 287
with refraction and local earthquake analysis [Smith 288
et al., 1982; Sparlin *et al.*, 1982; Sheriff and 289

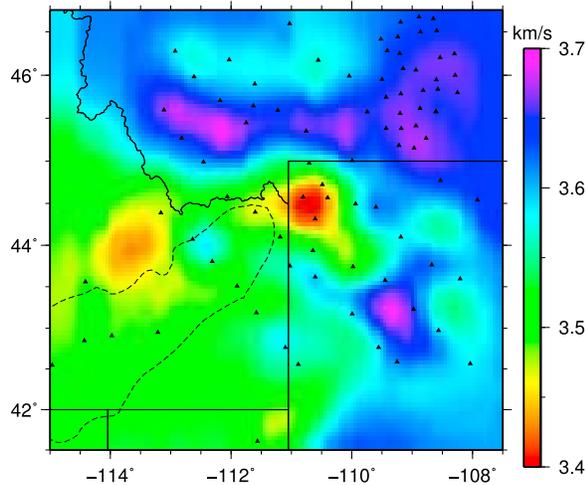


Figure 4. Mean crustal shear velocity from combined inversion of diffusive and ballistic surface wave dispersion measurements [Stachnik *et al.*, 2008]. PASSCAL stations used in this surface wave analysis are plotted as triangles.

290 *Stickney*, 1984; *Henstock et al.*, 1998; *Zeiler et al.*,
291 2005].

292 3.2. Common Conversion Point Images

293 [13] The CCP images (Figures 8–11) find that the
294 direct Moho arrival is well imaged by migration of
295 our radial receiver function data set. The ESRP/
296 Yellowstone Plateau parallel cross section (Figure 8,
297 cross section A) shows the Moho thickening to the
298 NE toward the Yellowstone Plateau and thick crust
299 within the Wyoming province. At the eastern end
300 of this cross section, the top of the Wyoming
301 Province high-velocity lower crustal layer is im-
302 aged at 35 km depth with the Moho at 52 km depth.
303 The NW–SE cross section through the Yellowstone
304 caldera (Figure 8, cross section B) shows a sharp
305 change in crustal thickness at the NE corner of
306 Yellowstone Park from thin crust beneath the
307 Montana Basin and Range to thick crust beneath
308 the Yellowstone caldera and the Wyoming province.
309 The ESRP perpendicular cross section across the
310 Picabo caldera (Figure 8, cross section C) shows
311 thick Wyoming province crust at the SE end of the
312 image and thin crust beneath the Montana Basin
313 and Range province. The ESRP crust beneath the
314 Picabo caldera is found to be seismically trans-
315 parent with upper crustal structure outside the
316 ESRP being truncated at the ESRP margins. In
317 general, the ESRP Moho is depressed by 2–4 km
318 with respect to the adjacent NW and SE margins
319 (see also Figure 7).

[14] The E–W Montana/northern Idaho cross sec- 320
tion (Figure 8, cross section D) shows the greatest 321
crustal thickness variation from 52 km beneath 322
eastern Montana to 35 km near the Eocene age 323
Bitterroot detachment and granitic batholith in 324
northern Idaho [Foster *et al.*, 2001]. The top of 325
the high-velocity lower crustal layer is also found 326
at the east end of this cross section. The N–S 327
Wyoming/Montana cross section (Figure 8, cross 328
section E) shows the thick (48–52 km) crust be- 329
neath the region shortening during the Laramide 330
orogeny and not affected by late Cenozoic ex- 331
tension [Dickinson, 2004]. The top of the high- 332
velocity lower crustal layer is imaged north of 333
44.5° latitude beneath the Billings array (Figure 1). 334
Two cross sections through the Billings array 335
(Figure 9) show the direct Moho arrivals from the 336
top of the high-velocity lower crustal layer at 28– 337
34 km depth. The thickness of this lower crustal 338
layer is up to 17 km thick with the layer thinning 339
to zero thickness at the NE and SW end of cross 340
section B. 341

[15] A final notable feature in the CCP images is a 342
paired positive and negative amplitude arrival at 12 343
and 18 km depth beneath the Yellowstone caldera 344
(Figures 8 (cross sections A and B) 10). This paired 345
arrival would be consistent with a low-velocity 346

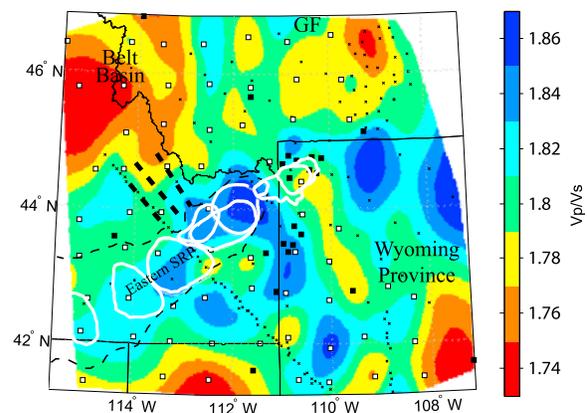


Figure 5. Crustal V_p/V_s map from interpolation of single-station H-K analysis results (i.e., Figure 3). The calderas are outlined with white solid lines, and the eastern Snake River Plain is outlined with the black dashed line. Stations are denoted as follows: transportable array stations (squares), PASSCAL experiment stations (dots), and University of Utah/Teton/National Seismic network stations (filled squares). The three dashed black lines to the NW of the Heise caldera field approximately locate the Beaverhead, Lemhi, and Lost River normal faults.

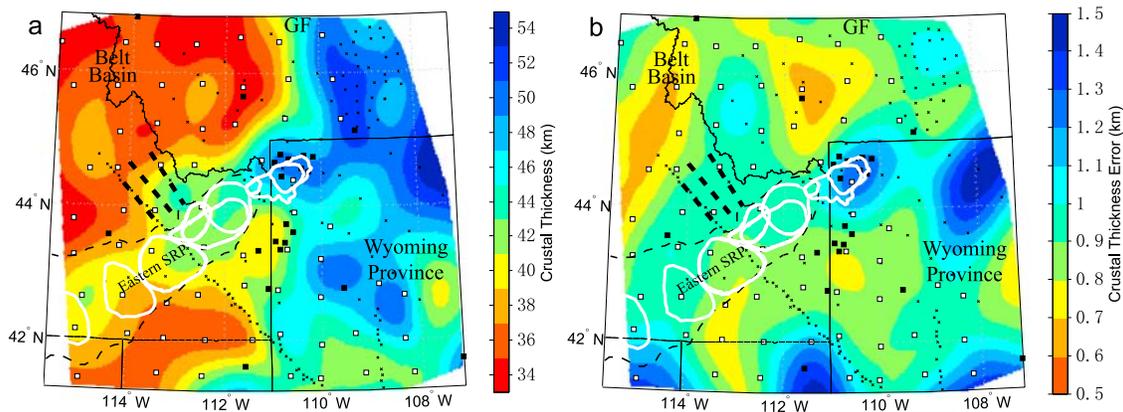


Figure 6. Crustal thickness and error map. (a) The calderas are outlined with white solid lines, and the eastern Snake River Plain is outlined with the black dashed line. Stations are denoted as follows: transportable array stations (squares), PASSCAL stations (dots), and University of Utah/Teton/National Seismic network stations (filled squares). (b) Crustal thickness standard deviation errors estimated via bootstrapping of the H-K analysis. This error analysis ignores velocity migration errors associated with the mean crustal shear velocity model (Figure 4) and the V_p/V_s map (Figure 5). The velocity migration uncertainties are assessed in section 2. The calderas are outlined with white solid lines, and the eastern Snake River Plain is outlined with the black dotted line.

347 zone: the 12 km negative polarity arrival mani-
348 festing a negative velocity gradient and the 18 km
349 positive polarity arrival manifesting a positive ve-
350 locity gradient. This paired arrival is directly under
351 the two most volcanically active regions of the

Yellowstone caldera between the Mallard Lake and 352
Sour creek resurgent Rhyolitic domes [Lowenstern 353
and Hurwitz, 2008]. A similar finding of an upper 354
crustal low-velocity zone is found by waveform 355
modeling of teleseismic S wave data from three 356

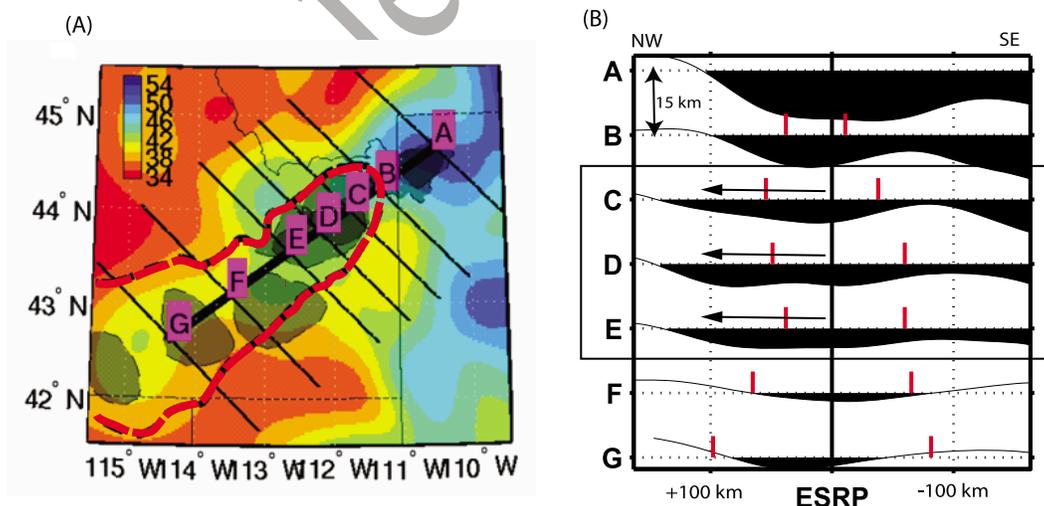


Figure 7. Eastern Snake River Plain crustal thickness maps and graphs. (a) The crustal thickness is indicated by the color bar. The cross sections shown in Figure 7b are labeled A–G, and the dark line down the center of the ESRP shows the zero values for the x axis coordinates in Figure 7b. The calderas are outlined as the gray shaded areas, and the eastern Snake River Plain is outlined with the black dashed line. (b) Crustal thickness graphs for cross sections A–G. The edge of the ESRP is marked by short red lines. The zero line for each graph is at 38 km depth, and a 15 km thickness variation is indicated by the scale bar. Cross sections C–E are boxed to indicate where the crust to the NW of the ESRP is anomalously thick with respect to cross sections F and G. The NW directed arrows indicated the direction of proposed lower crustal flow from beneath the Heise caldera field.

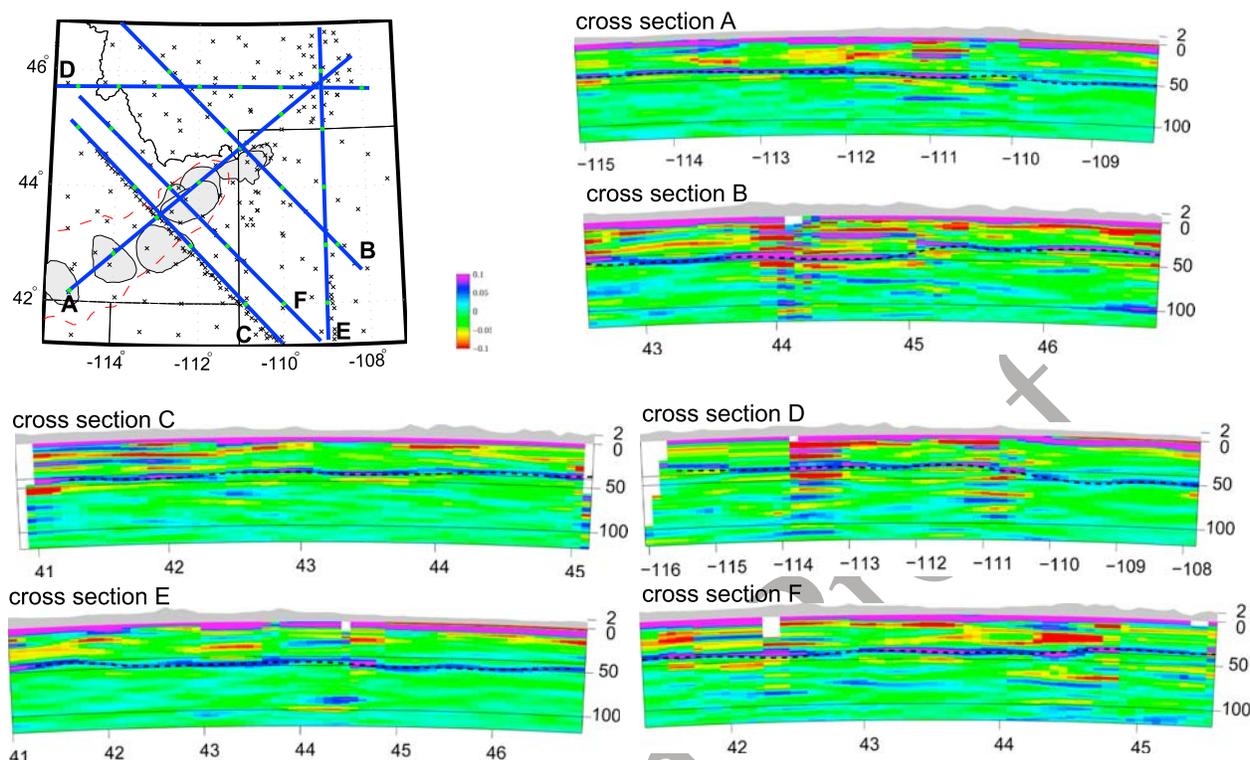


Figure 8. Common conversion point images. Topography is indicated as gray shading at the top of each section. Moho is indicated as black dotted line. Scale bar shows amplitude of positive (blue) and negative polarity (red) arrivals. Cross section A is the ESRP parallel section. Cross section B is the ESRP perpendicular section through Yellowstone Caldera. Cross section C is the ESRP perpendicular section through the Picabo caldera. Cross section D is the E–W section through Montana and northern Idaho. Cross section E is the N–S Wyoming province section. Cross section F is the ESRP perpendicular section through the Heise caldera field.

357 Yellowstone Park broadband stations [Chu *et al.*,
358 2009].

359 4. Discussion

360 4.1. Lower Crustal ESRP Outflow

361 [16] Our most important new result is our new
362 crustal thickness map which provides a synoptic
363 scale sufficient to assess the mass balances asso-
364 ciated with the magmatic inflation of the ESRP
365 crust and potential lower crustal outflow forced by
366 the midcrustal densification caused by this mag-
367 matic load. If the hypothesis that magmatic injec-
368 tion along the ESRP has stimulated crustal flow is
369 correct, then the time history of the crustal flow
370 should be correlated with the magmatic injection
371 history and the postcaldera lower crustal thermal
372 evolution which controls viscosity. Given the high
373 (>90 mW/m²) heat flow along the ESRP and its
374 margins [Blackwell and Richards, 2004] and the
375 finding of low flexural rigidity [Lowry and Smith,

1995; Lowry *et al.*, 2000], it seems plausible that
376 the lower crust is capable of flow driven by the
377 midcrustal sill load. Estimates of lower crustal flow
378 rates of 1–7 cm/yr (10–70 km/Myr) in regions of
379 elevated heat flow have been proposed where suf-
380 ficient pressure gradients in the lower crust exist
381 [Buck, 1991; Royden *et al.*, 1997, 2008]. Dividing
382 the 50–80 km lateral extent of the anomalously
383 thick crust to the NW of the Heise caldera field by
384 the mean age of the Heise caldera field (5 Ma)
385 provides a maximal lower crustal flow rate of about
386 1 cm/yr (10 km/Myr). 387

[17] The simplest crustal thickness evolution sce-
388 nario for the ESRP and its margins assumes that
389 prior to formation of the hot spot track, the crustal
390 thickness and density structure was uniform per-
391 pendicular to the ESRP between the Twin Falls and
392 Heise calderas and its adjacent margins. In addition,
393 post-hot spot track integrated crustal dilatation is
394 also assumed to be relatively uniform perpendicular
395 to the ESRP. Estimates of net dilatation at the NW
396 and SE ESRP margins is estimated as 15% and
397

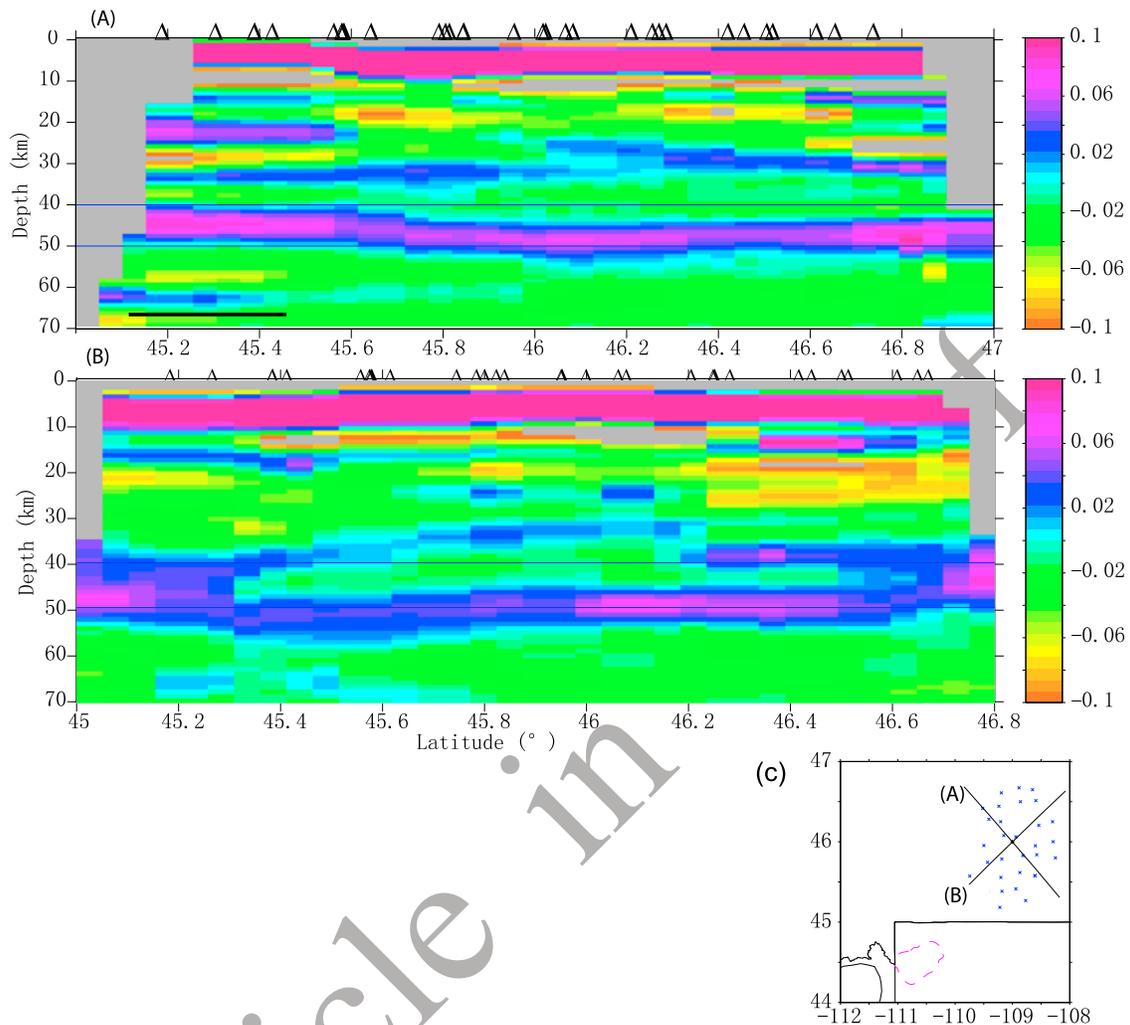


Figure 9. Billings array area common conversion point images. (a) Cross section A. (b) Cross section B. Note the positive arrival (blue-purple) that is found at 0–18 km above the Moho arrival at 47–52 km depth. (c) The locations of the cross sections and stations.

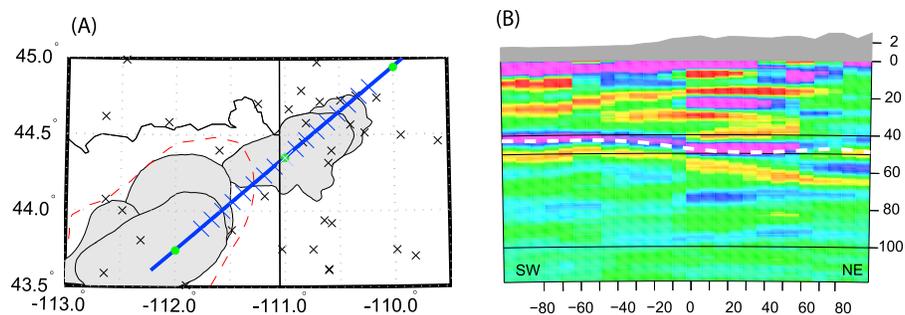


Figure 10. ESRP parallel common conversion point image highlighting structure under Yellowstone Caldera. (a) Cross section with calderas gray shaded and the ESRP outlined by the red dotted line. (b) Common conversion point image. Topography is indicated as gray shading at top of image, and the Moho is indicated as white dotted line.

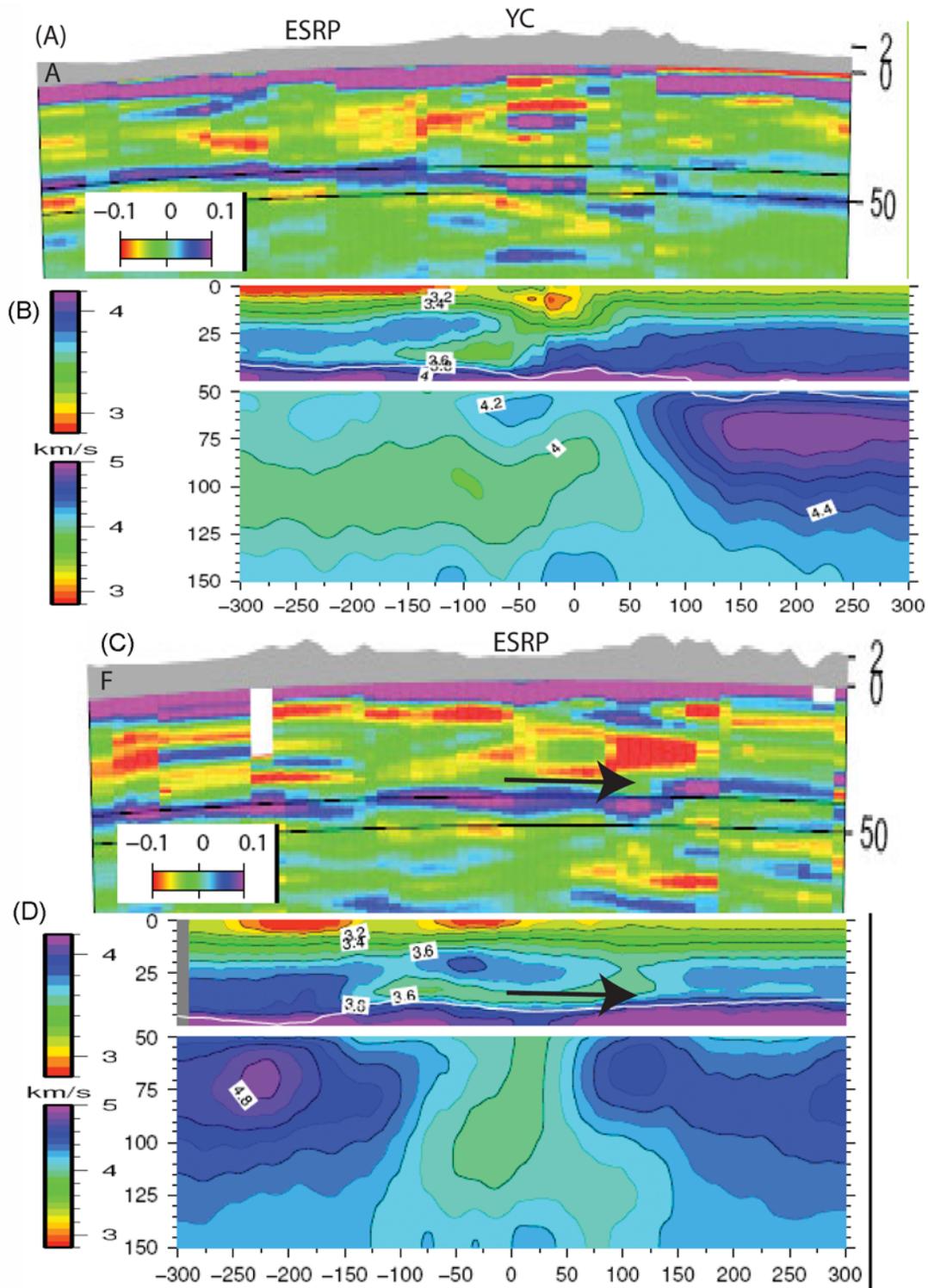


Figure 11. Coincident shear velocity tomogram [Stachnik *et al.*, 2008] and common conversion point images. The locations of cross sections A and F are shown in Figure 8. The gray shading at the top of each subplot is the elevation. The color scale for the CCP images denotes positive (blue) and negative polarity (red) arrivals. The color scale for the shear velocity is shown for the crust and mantle panels and the white line in the velocity images is the Moho. (a) CCP cross section A and (b) shear velocity tomogram. (c) CCP cross section F and (d) shear velocity tomogram. The arrows indicate the inferred dominant direction of lower crustal flow from beneath the Heise caldera field into the lower crust to the NW of the ESRP.



398 25% [Rodgers *et al.*, 2002]. Thus, to first order, we
399 assume that the ESRP initial conditions and inte-
400 grated dilatation perpendicular to the ESRP have
401 been relatively uniform. Given these assumptions,
402 the only change in the ESRP crustal thickness with
403 respect to its margins would result from mass ad-
404 ditions due to magmatism and sedimentation or
405 mass subtractions associated with caldera eruption
406 atmospheric plumes. The sedimentation addition is
407 estimated as <1 km and the atmospheric losses as
408 <0.75 km [Rodgers and McCurry, 2009]. These
409 two effects are opposite in sign and nearly cancel;
410 thus, the dominant ESRP mass variable is the cal-
411 dera forming crustal magmatic injections. In sum-
412 mary, the expected thickening of the ESRP crust
413 with respect to its margins is dictated by the pet-
414 rologic and caldera heat budget constraints that
415 suggest 8–14 km of mantle derived basaltic mag-
416 mas are required to fuel the calderas [Hanan *et al.*,
417 2008; McCurry and Rodger, 2009; Leeman *et al.*,
418 2008].

419 [18] The shear velocity tomogram finds a high-
420 velocity (>3.6 km/s contour) midcrustal layer
421 within the ESRP that starts at the NE end of the
422 ESRP beneath the Island Park caldera and extends
423 to the edge of our surface wave velocity sampling
424 at the Picabo caldera (Figures 11a and 11b). Be-
425 neath the Heise caldera field, the 3.6 km/s shear
426 velocity contour that outlines this layer resides
427 between 18–32 km depth (14 km thick). At the SE
428 end of the ESRP, the top of this high-velocity
429 midcrustal layer deepens by about 5 km. This layer
430 is thought to manifest the midcrustal sill (MCS)
431 complex that forms the mafic magmatic “tanks”
432 from whence the Rhyolitic caldera magmas were
433 derived via fractionation, assimilation, and hy-
434 bridization. Beneath this MCS layer, a relatively
435 low velocity lower crustal “wedge” is found that is
436 thickest beneath the Island Park caldera and pin-
437 ches out at the SW edge of the Heise caldera field.

438 [19] The ESRP perpendicular cross section
439 (Figures 11c and 11d) shows two interesting fea-
440 tures: the MCS layer dips to the NW and the low-
441 velocity (<3.6 km/s) lower crust extends up to 50 km
442 to either side of the ESRP margins. With respect to
443 the mantle velocity anomalies shown in the tomo-
444 grams, the ESRP-YP physiographic province is
445 underlain by very low (4 km/s) shear velocities at
446 75–125 km depth with a relatively high velocity
447 mantle lid between the Moho and 75 km depth
448 (Figures 11a and 11b). The low ESRP mantle ve-
449 locity rapidly changes beneath the Beartooth
450 Mountains to cratonic lithospheric velocity values of
451 4.9 km/s [Artemieva, 2009; Bedle and van der Lee,

2009]. In the ESRP perpendicular cross section 452
(Figures 11c and 11d), the low-velocity ESRP 453
mantle is found directly beneath the 90 km wide 454
ESRP with higher-velocity mantle lithosphere be- 455
neath SW Wyoming (SE end) and western Montana 456
(NW end). 457

[20] Based on the above observations, we believe a 458
good circumstantial case is made for the flow of 459
magmatically thickened lower crust from beneath 460
the Heise caldera field into the adjacent NW margin 461
crust. However, this statement begs the question 462
of where the magmatically thickened ESRP crust 463
has flowed from beneath the older Twin Falls and 464
Picabo caldera fields. Two differences are noted 465
between the Heise caldera field and the older Twin 466
Falls and Picabo calderas. First, these two older 467
calderas are spaced farther apart along the ESRP 468
(Figure 1) indicating less mass flux into the ESRP 469
per unit area. Second, these caldera fields have had 470
more integrated dilatation with respect to the Heise 471
caldera field because extension associated with 472
caldera formation began earlier in this region 473
[Anders *et al.*, 1989; Pierce and Morgan, 1992]. In 474
addition, the proximity of these older calderas to the 475
concentrated extension of the western Snake River 476
Plain graben [Cummings *et al.*, 2000] and the 477
Northern Nevada Rift [Glen and Ponce, 2002] is 478
noted; these regions of concentrated extension could 479
create lateral pressure gradients in the lower crust 480
that would promote the flow of lower crustal mass 481
from the magmatically thickened calderas areas. 482

4.2. Yellowstone Caldera Low-Velocity Zone 483

[21] Modern day gas fluxes near recent intracaldera 485
basaltic eruptions [Lowenstern and Hurwitz, 2008] 486
and deformation monitoring [Chang *et al.*, 2007; 487
Puskas *et al.*, 2007] suggest that ongoing post 488
0.6 Ma Yellowstone caldera magmatic activity is 489
occurring. The paired negative/positive amplitude 490
arrivals at 12–18 km depth found beneath the Yel- 491
lowstone caldera by the CCP images (Figures 8 (cross 492
sections A and B) and 10) are consistent with other 493
geophysical data that suggest the Yellowstone cal- 494
dera low-velocity and low-density anomalies extend 495
to about 20 km depth: the low velocities beneath the 496
caldera in our shear wave tomogram (Figures 11a 497
and 11b) and gravity modeling [DeNosaquo *et al.*, 498
2010]. A tomogram constructed from measured 499
local earthquake traveltimes finds a low (5.4 km/s) 500
P wave velocity at 8 km depth, but cannot resolve 501
structure below 12 km depth [Husen *et al.*, 2004]. 502
Waveform modeling of S-P precursors from three 503

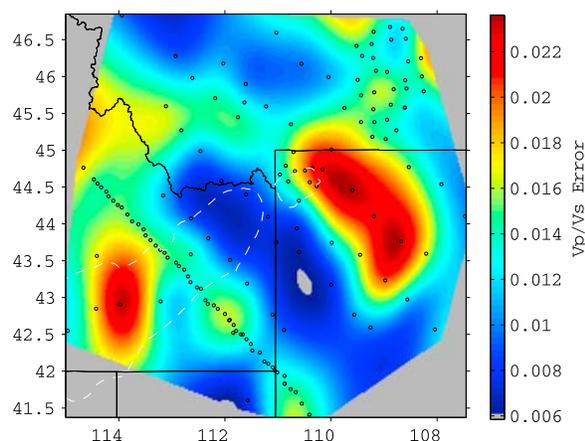


Figure A1. V_p/V_s map of standard deviation. The standard deviation was calculated via bootstrap resampling of the H-K analysis to calculate an ensemble of V_p/V_s and crustal thickness values.

504 broadband stations near the Mallard Lake dome
505 requires a substantial low-velocity zone at 5–8 km
506 with a thickness of 3–5 km [Chu et al., 2009].
507 Thus, our finding of a 6 km thick low-velocity zone
508 with its top at 12 km depth provides a depth range
509 over which basaltic magma is being staged beneath
510 the Yellowstone caldera. In the near future, this
511 finding can be tested via higher-resolution ambient
512 noise surface wave imaging with the deployment of
513 new broadband seismometers within Yellowstone
514 Park in 2010.

515 Appendix A

516 [22] Map of the standard deviation of V_p/V_s mea-
517 surements (Figure A1) finds an average of 0.015.
518 The V_p/V_s errors were estimated for each station
519 via bootstrap resampling of the peak amplitude of
520 the H-K image. These individual station standard
521 deviation estimates are interpolated using a two-
522 dimensional spline that fits the individual station
523 error estimates to within their individual error bars.

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